

Lols update, perspectives for low energy and reaccelerated beams

Pierre Delahaye



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Perspectives with low energy beams

Some perspectives at LIRAT, but limited due to beam purity issues



At longer term, low energy physics @ DESIR

- Upgraded SPIRAL 1 facility
- S3 – Low Energy Branch **when available**



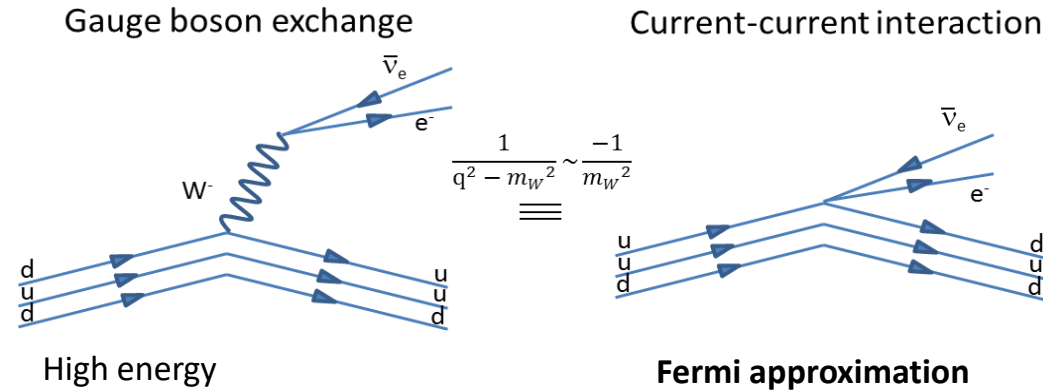
DESIR

Unique perspectives for SPIRAL 1 beams at DESIR

${}^6\text{He}$ searching for

- * dark decay (H. Savajols DESIR session)
- * evidence for tensor currents (bSTILED)

Nuclear β -decay as a laboratory for weak interaction



$$\mathcal{H} = \frac{G_F V_{ud}}{\sqrt{2}} \sum_{i=V,A,S,T,P} (\bar{\Psi}_p O_i \Psi_n) [\bar{\Psi}_e O^i (C_i + C'_i \gamma_5) \Psi_\nu] + h.c.$$

Accessible via **Ft-values**
 $\sim 2 \cdot 10^{-4}$ thanks to $0^+ \rightarrow 0^+$ transitions

Hardy and Towner, arXiv:1807.01146v1

T. D. Lee and C. N. Yang, Phys. Rev. 104, 254 (1956).

Explicit parity violation $C'_i \neq 0$

Sum of Lorentz invariants

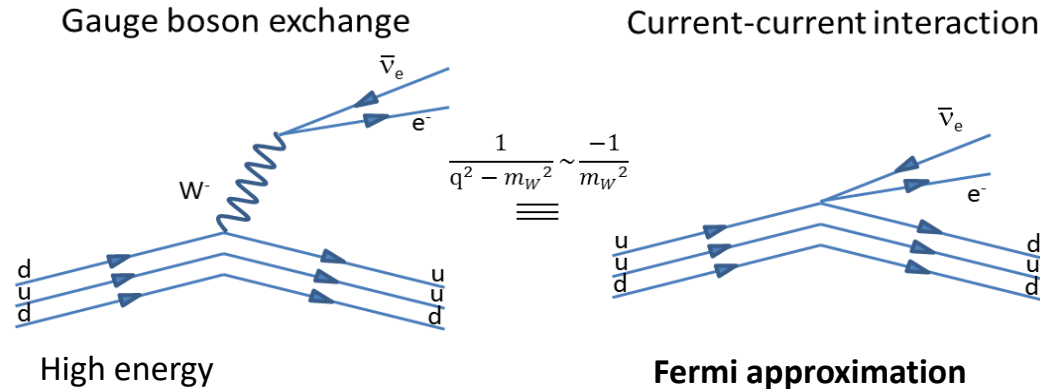
$|V_{ud}| = 0.97370(10)_{\text{NP}}(10)_{\text{RC}}$ Seng et al., PRL 121 (2018)

$|V_{ud}| = 0.97389(10)_{\text{NP}}(14)_{\text{RC}}$ Czarnecki, Marciano, Sirlin, Phys Rev D 100 (2019)

Belfatto et al., EPJC (2020)

CKM: is deviation from unitarity a sign of new physics at the TeV scale?

Nuclear β -decay as a laboratory for weak interaction



$$\mathcal{H} = \frac{G_F V_{ud}}{\sqrt{2}} \sum_{i=V,A,S,T,P} (\bar{\Psi}_p O_i \Psi_n) [\bar{\Psi}_e O^i (C_i + C'_i \gamma_5) \Psi_\nu] + h.c.$$

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Explicit parity violation $C'_i \neq 0$

Sum of Lorentz invariants

Accessible via **correlation measurements**

Standard Model

At quark level: $C_V = -C_A = 1, C_S = C_T = 0$

And $C_i = -C'_i$ for **maximal parity violation**

Real for T reversal/CP conserving interactions

Perspectives with low energy beams

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DESIR

Unique perspectives for SPIRAL 1 beams at DESIR

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Perspectives with low energy beams

Some perspectives at LIRAT, but limited due to beam purity issues

At longer term, low energy physics @ DESIR

- Upgraded SPIRAL 1 facility
- S3 – Low Energy Branch **when available**



DESIR

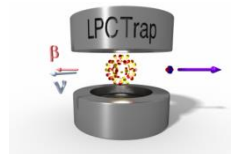
Unique perspectives for SPIRAL 1 beams at DESIR

* FT-value measurements

Extension of $0^+ \rightarrow 0^+$ superallowed to mirror transitions for the determination of V_{ud}

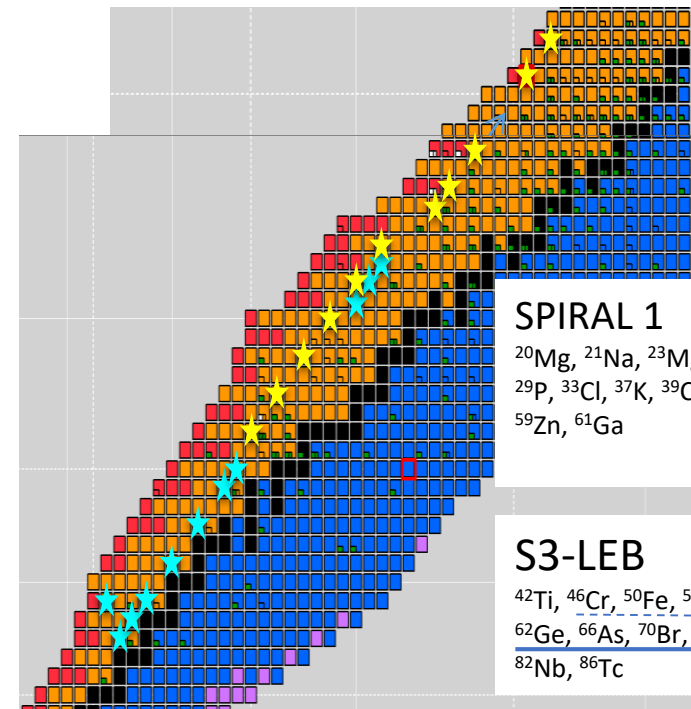
* Correlation measurements

$a_{\beta V}(\rho)$ in the mirror decay of ^{21}Na , ^{23}Mg , ^{33}Cl , ^{37}K , ^{39}Ca , ^{41}Sc



Test of CP violation in beta-decay
 $^{23}\text{Mg}^+$, $^{39}\text{Ca}^+$ beams from SPIRAL 1

The MORA project



SPIRAL 1

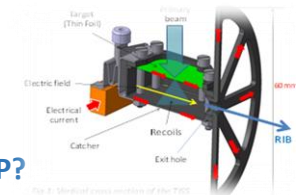
^{20}Mg , ^{21}Na , ^{23}Mg , ^{25}Al ,
 ^{29}P , ^{33}Cl , ^{37}K , ^{39}Ca , ^{57}Cu ,
 ^{59}Zn , ^{61}Ga

S3-LEB

^{42}Ti , ^{46}Cr , ^{50}Fe , ^{54}Ni , ^{58}Zn ,
 ^{62}Ge , ^{66}As , ^{70}Br , ^{74}Kr ,
 ^{82}Nb , ^{86}Tc

* **Availability, operationability, for producing scientific highlights**

----- 1



TULIP?

+ Mass measurements, 2p radioactivity with ^{22}Al , ^{31}Ar , ^{26}P , ^{35}Ca etc.

What can be done before DESIR?

- **Preparing DESIR!**
 - N=Z nuclei with TULIP
 - ? Mass measurements and identification with PILGRIM at LIRAT
 - MR TOF MS: a **must-have** for beam identification at RIB facilities
 - Unique tool for long-lived beams (^{56}Ni , ^{48}Cr etc) yield measurement
 - Unique tool for identifying contaminants
 - Powerful/cost effective tool for purifying beams for beta decay experiments
 - Feasibility to be assessed: Human resources, Planification for DESIR, need for RFQ cooler
 - **Do we just wait that DESIR is built?**
 - Lol Pauline Ascher et al, mass measurements for probing
 - N=28 island of inversion (isotopic chains up to ^{49}Ar , ^{48}Cl , ^{43}P)
 - Vpn and possibility for 2n decay in $^{24-28}\text{F}$ and $^{22-24}\text{O}$
 - ^{23}Mg and ^{39}Ca beam development for MORA

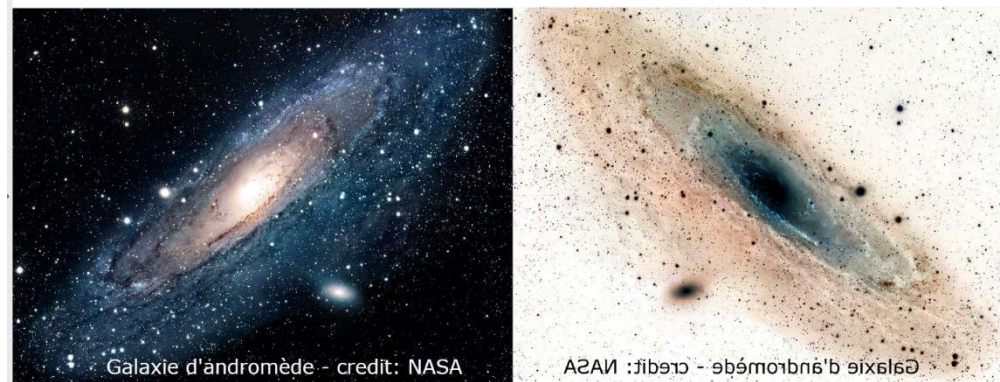
Search for new physics via the D correlation measurement

$$D \frac{\langle \vec{J} \rangle}{J} \cdot \left(\frac{\vec{p}_e}{E_e} \times \frac{\vec{p}_\nu}{E_\nu} \right)$$

T reversal odd

A non-zero D can arise from CP violation

- CP violation observed in the K, B and D - meson decays is not enough to account for the large matter – antimatter asymmetry
- T-odd correlations in beta decay (D and R) and n-EDM searches are sensitives to larger CP violations by 5 to 10 orders of magnitude



See P. Herczeg, Prog. Part. Nucl. Phys. 46 (2001) 413.

Below 10^{-4} , Final State Interactions mimic a non zero correlation

Best measurement so far: $D_n = (-0.94 \pm 1.89 \pm 0.97) \cdot 10^{-4}$ *emiT collaboration, PRL 107, 102301 (2011)*,
 $D_{19Ne} = (1 \pm 6) \cdot 10^{-4}$ *Calaprice, Hyp. Int. 22(1985)83*

MORA: D correlation measurement to the 10^{-5} level with some beam, laser and trapping R&D

- Search for leptiquarks beyond the TeV scale
- First approach to probe of Final State interaction effects

MORA: Best candidates for D measurement

$$D \equiv \sin(\varphi_{AV}) \cdot \underbrace{\frac{2\rho}{1+\rho^2} \cdot \left(\frac{J}{J+1}\right)^{1/2}}_{F(X)}$$

Sensitivity to CP violating phase between V and A currents

Search for New Physics

- **Direct constraints** on CP-violating Wilson coefficients in the nucleon-level EFT
- via **the L-R symmetric model**
- via **the LQ model**

Neutron and mirror nuclei (N=Z-1): strong mixed (GT+ Fermi) transitions between analog states

	n	¹⁹ Ne	²³ Mg	³⁵ Ar	³⁹ Ca
Sensitivity $F(X)$	0,43	-0,52	-0,65	0,41	0,71
D_1 ($\times 10^{-4}$)	0,108	2,326	1,904	0,386	-0,489
D_2 ($\times 10^{-4}$)	0,023	0,169	0,099	0,010	-0,024

$$D_n = (-0.94 \pm 1.89 \pm 0.97) \cdot 10^{-4} \quad D_{^{19}\text{Ne}} = (1 \pm 6) \cdot 10^{-4}$$

Best measurement so far, *statistics limited*

$$D_{FSI}(p_e) = \left(D_1 \cdot \frac{p_e}{p_{e\max}} + D_2 \cdot \frac{p_{e\max}}{p_e} \right) \times 10^{-4}$$

Callan and Treiman, Phys. Rev. 162(1967)1494.
Chen, Phys. Rev. 185(1969)2003.

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MORA: Alkali earth elements for in trap laser ion polarization

1st candidate; 10^5 pps from JYFL 2nd candidate, R&D for ISOL
 $>10^8$ pps from SPIRAL 1 production required

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10^7 pps requested

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Prospects with ^{39}Ca

^{39}Ca : a counter part of ^{23}Mg for looking for New Physics and verifying the Final State Interactions

Place and type of measurement	Trapped ions /cycle	Decays/s	Beam time (days)	Detected coinc. (P)	σ_p stat (%)	Detected coinc. (D)	σ_D stat*
JYFL: P - ^{23}Mg	2,00E+04	1,23E+03	8	1,7E+05	1,9E+00	1,5E+06	1,0E-03
JYFL: D - ^{23}Mg	2,00E+04	1,23E+03	32	6,7E+05	9,4E-01	6,1E+06	5,2E-04
DESIR: D - ^{23}Mg	5,00E+06	3,07E+05	24	1,3E+08	6,9E-02	1,2E+09	3,8E-05
DESIR: D - $^{39}\text{Ca}^{**}$	5,00E+06	4,03E+06	24	1,7E+09	1,5E-03	1,5E+10	1,0E-05

* All measurements so far have been statistics dominated

** Assuming **10⁷ pps produced** from SPIRAL 1 or S3 – LEB

4 years for developping ^{39}Ca at GANIL/SPIRAL2, in parallel of

- First *D* measurement(s) and careful study of systematic effects at JYFL
- Theoretical studies at IJCLab - [Adam Falkowski](#), [Antonio Rodriguez-Sanchez](#)
 - Full classification of models of New Physics probed by a $D \sim 10^{-5}$ measurement
 - Verification of Final State Interactions calculations with modern EFTs
- Construction of DESIR and associated instruments (HRS, Lasers cabins, ...)

Prospects with ^{39}Ca

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JYFL: P - ^{23}Mg	2,00E+04	1,23E+03	8	1,7E+05	1,9E+00	1,5E+06	1,0E-03
JYFL	The ultimate sensitivity of MORa to New Physics technically depends on: <ul style="list-style-type: none"> • Statistics • Control of systematic effects Taking example of successful measurements carried out with neutron decay, and EDM A regular access to online run with possible long data accumulation periods* is highly desired * 3 to 4 weeks typically						4
DESIR							5
DESIR							5

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LoI6_20 @ GANIL: «Li-Alpha» -Measurement of ${}^8\text{Li}(\alpha, n){}^{11}\text{B}$ cross section with ACTAR

See T. Roger's presentation on ACTAR

${}^8\text{Li}$ beam development @ SPIRAL1

New Proposal for December 2022 GANIL PAC - Spokesperson: M.G. Pellegriti, I. Lombardo, T. Roger

ASTROPHYSICS: nucleosynthesis in the early evolutionary stage of our universe, starting of r-process nucleosynthesis in supernovae, collapsars and neutron star mergers

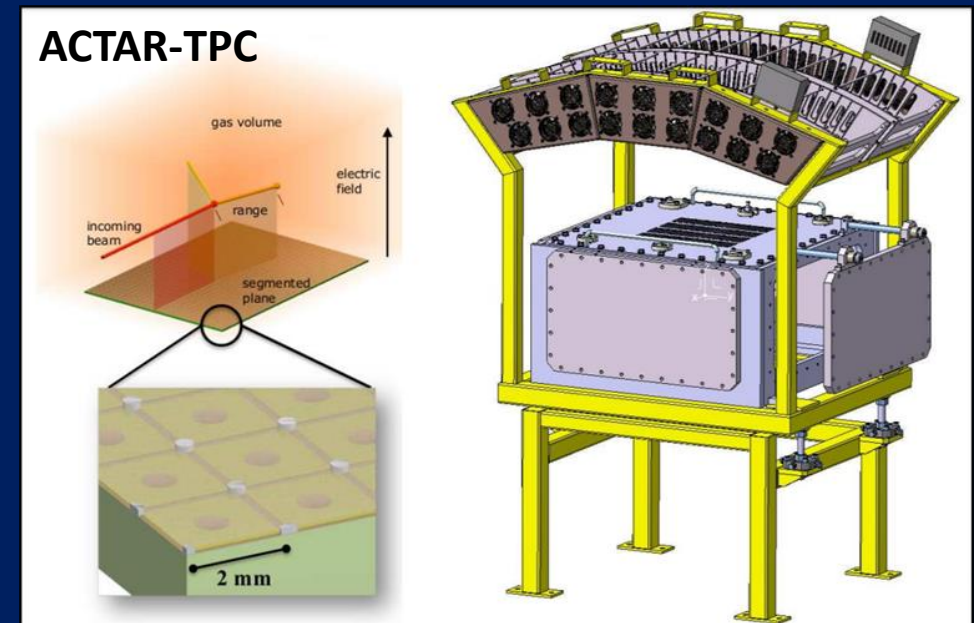
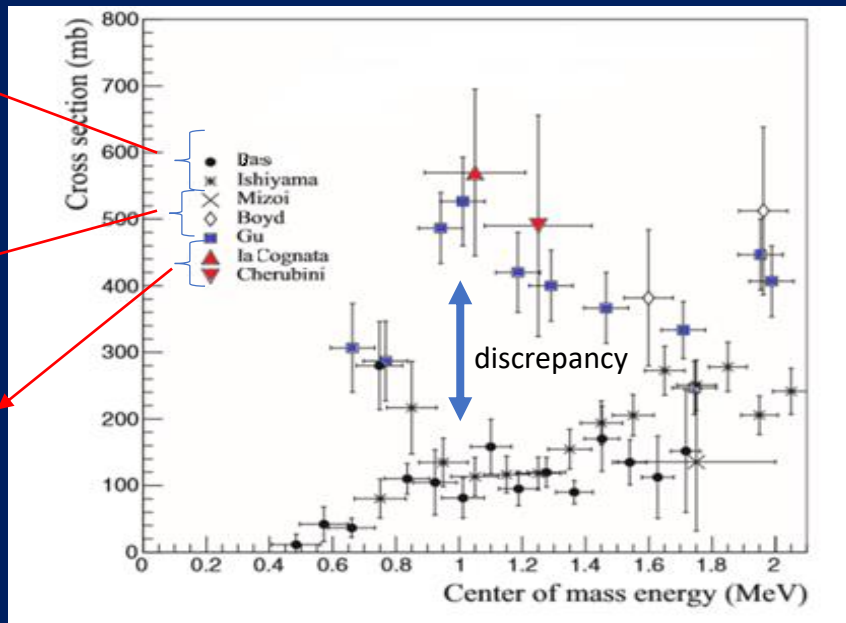
STRUCTURE: α -structure of ${}^{12}\text{B}$

inclusive ${}^{11}\text{B}$ detection with a factor 5 smaller pitch

EXCLUSIVE : $n + {}^{11}\text{B}$

INCLUSIVE: ${}^{11}\text{B}$

INCLUSIVE: n



Huge discrepancy between inclusive and exclusive measurements !!

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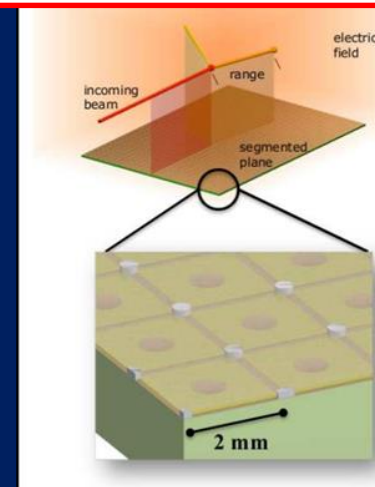
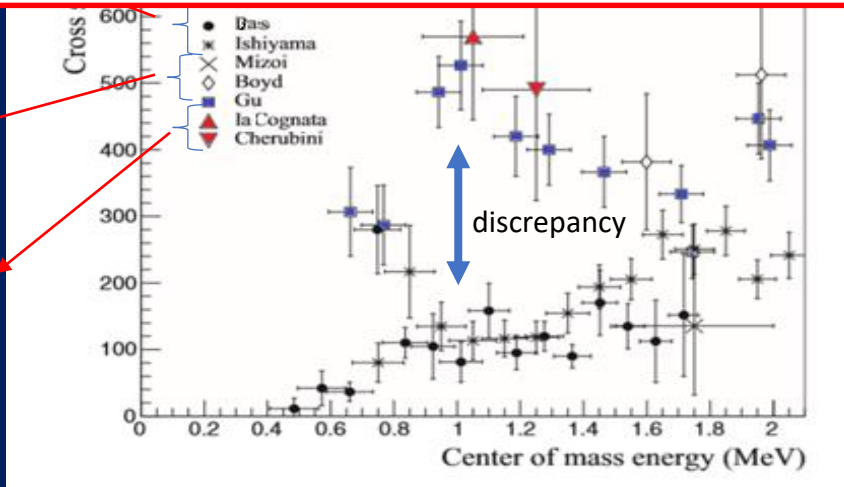
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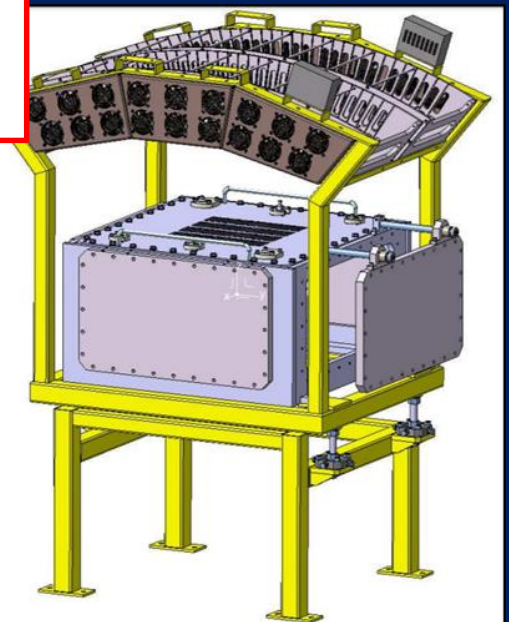
INCLUSIVE: n

Tests are completed:

- Beam produced with FEBIAD source, 10^6 pps feasible with ${}^{13}\text{C}$
- 2% for reacceleration **through CIME at 1.2 AMeV, no charge breeding**
- Interesting prospects for ${}^9\text{Li}$ if Mononake instead of FEBIAD



with a factor 5 smaller pitch



Huge discrepancy between inclusive and exclusive measurements !!

^{79}Se and ^{85}Kr for nuclear astrophysics using MUGAST

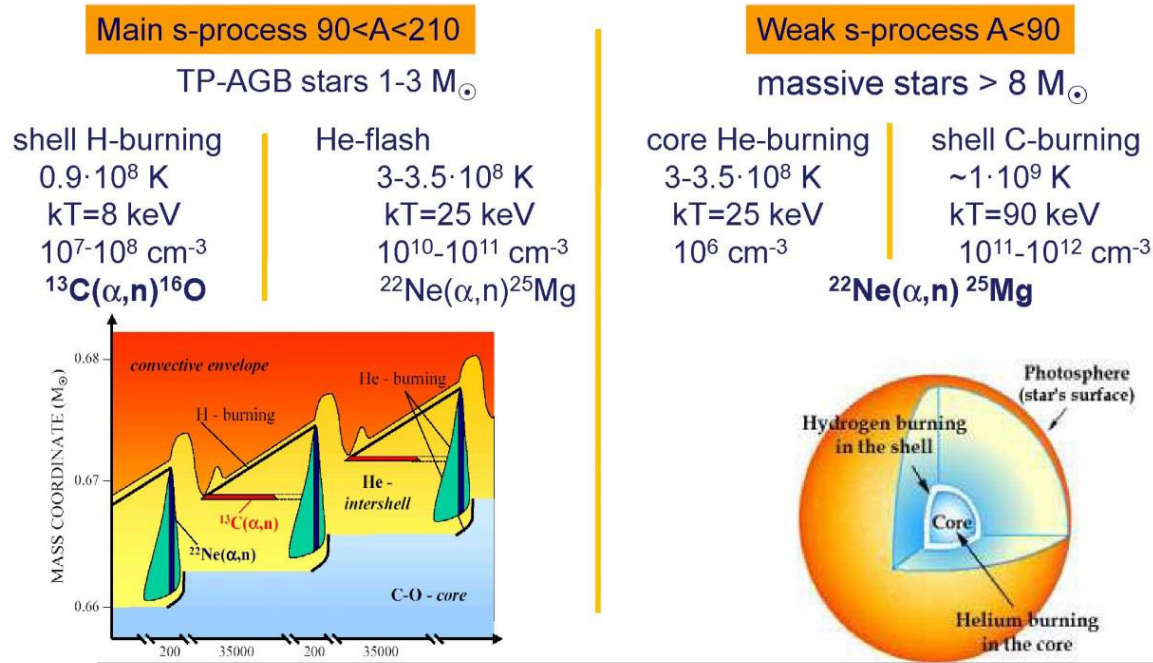
See O. Sorlin's presentation on LISE

Neutron capture at the ^{85}Kr s-process branching: F. Recchia and G. De Angelis (LoI/proposal 2018, resubmitted)

The $^{79}\text{Se}(n,\gamma)$ capture cross section via the $^{79}\text{Se}(d,p)^{80}\text{Se}$ reaction, G. De Angelis et al. (LoI 2019)

(p,d) surrogate reactions to (n, γ)

(n, γ) cross section for constraining s process path, associated neutron densities and temperature



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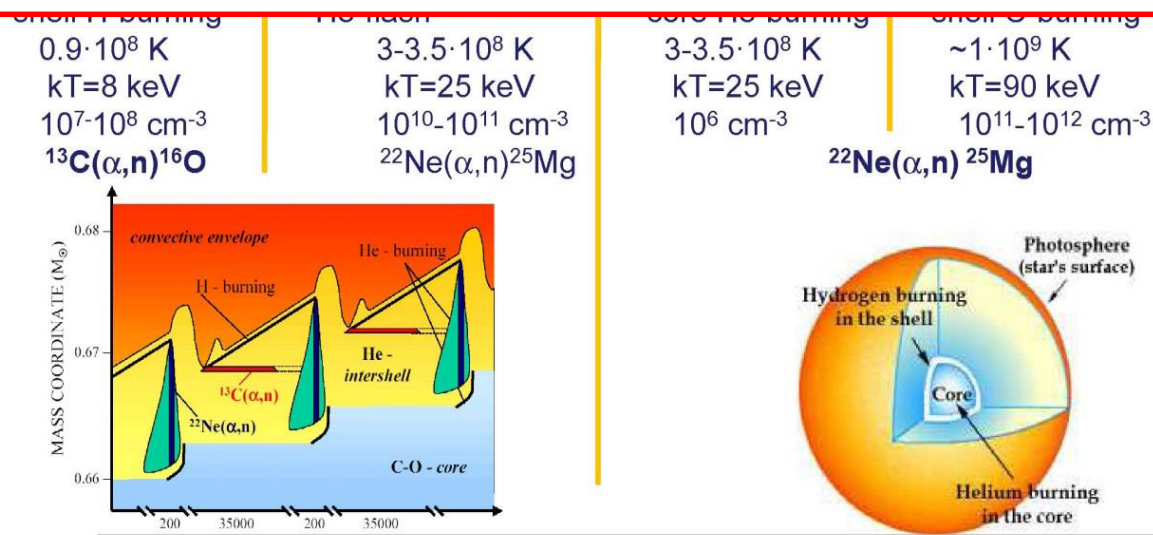
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(p,d) surrogate reactions to (n, γ)

(n, γ) Very long lived beams:

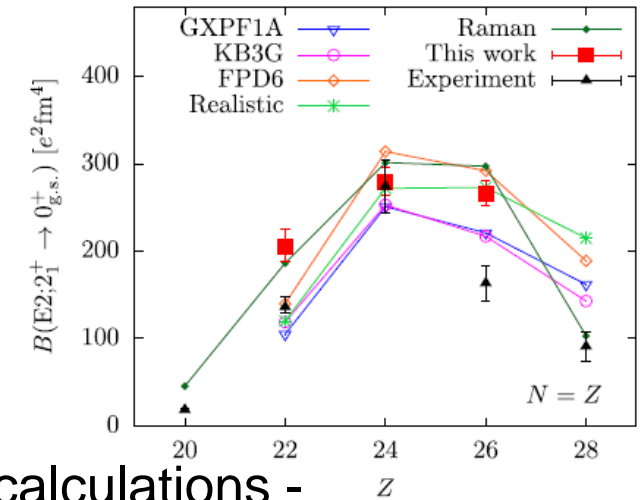
- ^{85}Kr (10.7y) **commercially available** for ORANO as a gas, 2 bottles (120MBq) purchased
- ^{79}Se (327ky) **not yet available** while ^{75}Se is. Latest test with ^{84}Kr @SPIRAL1: ^{71}Se produced.

Production at SPIRAL 1 with or without the CSS?



^{48}Cr physics cases

- Lol E. Clément, W. Korten 2019: probing **rotational character** of the most collective N=Z nucleus in fp shell via **safe Coulomb excitation**
 1. An independent measurement of the $B(E2, 4^+ \rightarrow 2^+)$
 2. The study of the non-yrast states collectivity
 3. A first measurement of the static quadrupole moment $Q_s(2^+)$
- ^{48}Cr using MUGAST (see presentation by O. Sorlin)
 - **Spectroscopic factors via (d,p) transfert** to compare with LSSM calculations - ANTOINE K3B interaction. A Gadea et al.
 - **Probing the effect of deformation for T=1 np pairing** in $^{48}\text{Cr}(^3\text{He}, \text{py})^{50}\text{Mn}$ reaction (cryogenic target). M. Assié et al.



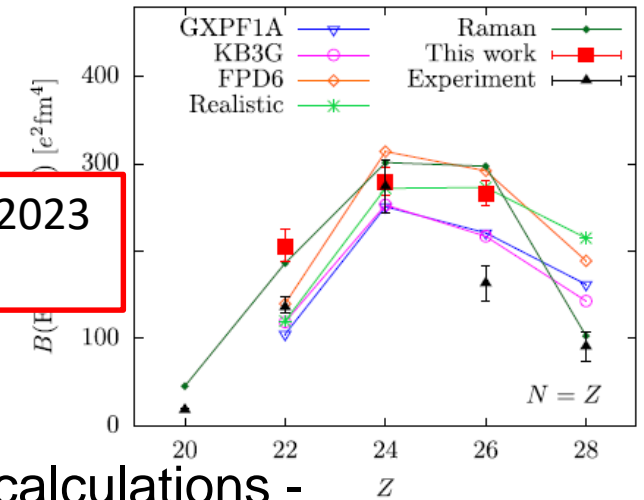
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3. A first measurement

Focus on ^{48}Cr , for which measurements would have to start in 2023
Production with the FEBIAD source at SPIRAL 1

- ^{48}Cr using MUGAST (see presentation by O. Sorlin)
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Other Physics cases using MUGAST

56 Ni 28 Ni 28 6.075 d 0+ M -53907.5 (0.4) β ⁺ =100%	57 Ni 28 Ni 29 35.60 h 3/2- M -56083.8 (0.6) β ⁺ =100%	58 Ni 28 Ni 30 stable 0+ M -60228.7 (0.4) Abundance=68.077 (19)% 2β ⁺ ?
55 Co 27 Co 28 17.53 h 7/2- M -54029.9 (0.4) β ⁺ =100%	56 Co 27 Co 29 77.236 d 4+ M -56040.4 (0.5) β ⁺ =100%	57 Co 27 Co 30 271.70 d 7/2- M -59345.6 (0.5) EC=100%
54 Fe 26 Fe 28 364 ns 10+ Ex 8527.1 (1.1) IT=100%	55 Fe 26 Fe 29 2.744 y 3/2- M -57481.3 (0.3) EC=100%	56 Fe 26 Fe 30 stable 0+ M -60607.1 (0.3) abundance=91.754 (36)%

MUGAST Prepac 2019

Astro	$^{55}\text{Co}, ^{57}\text{Ni}(d,p)$ $^{33}\text{Cl}(^3\text{He},d)^{34}\text{Ar}$ → <i>G. Lotay (Surrey)</i> Readily produced
Pairing	$^{56}\text{Ni}, ^{48}\text{Cr}(^3\text{He},p)$ <i>M. Assié, IPNO</i>

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Already several Lols on ^{56}Ni in 2018 (M. Assié / F. Flavigny)

Long lived beams, not easy to identify, probably quite polluted → stripping necessary

Production with the FEBIAD source at SPIRAL 1 to be discussed for possible scheduling

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Summary table for developments

Priority 0 (accepted experiments)						
Isotope	Intensity (min)	Energy	Timescale	Reference	Production method	Difficulty
85Kr	1e6 pps	10 AMeV	<1 year	Lol Recchia 2019 -		
Priority 1 (endorsed Lol or ANR)						
Isotope	Intensity (min)	Energy	Timescale	Reference	Production method	Difficulty
8Li	10 ⁴ -10 ⁵ pps	1.4 AMeV	2023-25	Lol Grazia 2020	Fragmentation 12C/13C	
48Cr	5 10 ⁴ pps	4.4 AMeV	2024	Lol Clément 2019		
39Ca	> 10 ⁷ pps	15-30 keV	2027	Lol Delahaye 2021	?	
23Mg	10 ⁸ pps	15-30 keV	2027	Lol Delahaye 2021		
74Rb		15-30 keV		ANR TULIP	Fusion evaporation	
79Se	10 ⁵ pps	9 AMeV	2023-25	Lol De Angelis 2019		
Priority 2: recent Lols, recent expressed interests						
Isotope	Intensity (min)	Energy	Timescale	Reference	Production method	Difficulty
56Ni	5 10 ⁴ pps	10-15AMeV	2023-25	Lols Flavigny Assié 2018		
55Co	~5E4 pps	~10 AMeV	2023-25	Mugast Prepac Lotay		
57Ni	~5E4 pps	~10 AMeV	2023-25	Mugast Prepac Lotay		
33Cl	~5E4 pps	~10 AMeV	2023-25	Mugast Prepac Lotay		
67As	~5E4 pps	~10 AMeV	2023-25	Mugast Prepac Mengoni		
118,120Xe	>1e3 pps	4.4 AMeV	2023-25	Lol Clément 2018	Fusion Evaporation	
70Se	>10 ⁴ pps	10-15AMeV	2023-25	Com. Valerian Alcindor		
69Cu	> 10 ⁴ pps	10-15AMeV	2023-25	Com. N. De Séreville		

Open for discussion

→ Originally planned on AGATA, difficult beam

→ TULIP?

Summary table for developments

Priority 3: long term opportunities, older Lols with case to be confirmed							
Isotope	Intensity (min)	Energy	Timescale	Reference	Production method	Difficulty	
41-48Cl	1pps	15-30keV	2027?	Lol Ascher 2016		None 74Kr	
39-43P	1pps	15-30keV	2027?	Lol Ascher 2016			
24-28F	1pps	15-30keV	2027?	Lol Ascher 2016			
46-49Ar	1pps	15-30keV	2027?	Lol Ascher 2016			
22-24O	1pps	15-30keV	2027?	Lol Ascher 2016			
25Al	1e5 pps	10-15AMeV		Lol De Séreville 2016	AIF		
30P	1e5 pps	10-15AMeV		Lol De Séreville 2016			
74Kr or 74Rb	1e3 pps	5AMeV?		Lol Singh 2016	78Kr fragmentation		
148Gd	1e4pps	<5AMeV?		Lol Caamaño 2016	Fusion Evaporation		
158Ho	1e4pps	<5AMeV?		Lol Caamaño 2016	Fusion Evaporation		
160Yb	1e4pps	<5AMeV?		Lol Caamaño 2016	Fusion Evaporation		
170Lu	1e4pps	<5AMeV?		Lol Caamaño 2016	Fusion Evaporation		
26mAl	1e6pps	3AMeV?		E792S			
Priority 3 (still 3): beams only discussed during the SPIRAL 1 WS in 2016							
45K	1e6 pps	2.5 AMeV		Prepac AGATA 2016			
25Na	1.8e6 pps	~2.5 AMeV		Prepac AGATA 2016			
21F	> 1e4pps?	3AMeV		Prepac AGATA 2016			
25Na	> 1e4pps?	3AMeV		Prepac AGATA 2016			
21,22, 27Mg	> 1e4pps?	3AMeV		Prepac AGATA 2016			
29Al	> 1e4pps?	3AMeV		Prepac AGATA 2016			
33P	> 1e4pps?	3AMeV		Prepac AGATA 2016			
35K	> 1e4pps?	3AMeV		Prepac AGATA 2016			
21,26Na	1e4 pps	15-30 keV	2027 DESIR	WS SPIRAL 1 2016			
22,23Mg	1e4 pps	15-30 keV	2027 DESIR	WS SPIRAL 1 2016			
29P	1e4 pps	15-30 keV	2027 DESIR	WS SPIRAL 1 2016			
37K	1e4 pps	15-30 keV	2027 DESIR	WS SPIRAL 1 2016			

Open for discussion